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Petersen, Peter and Wylie, William, "Rigidity of Gradient Ricci Solitons" (2007). *Mathematics - Faculty Scholarship*. 127.

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RIGIDITY OF GRADIENT RICCI SOLITONS

PETER PETERSEN AND WILLIAM WYLIE

ABSTRACT. We define a gradient Ricci soliton to be rigid if it is a flat bundle $N \times_{\Gamma} \mathbb{R}^k$ where N is Einstein. It is known that not all gradient solitons are rigid. Here we offer several natural conditions on the curvature that characterize rigid gradient solitons. Other related results on rigidity of Ricci solitons are also explained in the last section.

1. INTRODUCTION

A Ricci soliton is a Riemannian metric together with a vector field (M, g, X) that satisfies

$$\text{Ric} + \frac{1}{2}L_X g = \lambda g.$$

It is called shrinking when $\lambda > 0$, steady when $\lambda = 0$, and expanding when $\lambda < 0$. In case $X = \nabla f$ the equation can also be written as

$$\text{Ric} + \text{Hess} f = \lambda g$$

and is called a gradient (Ricci) soliton. We refer the reader to [5, 6, 7, 8] for background on Ricci solitons and their connection to the Ricci flow. It is also worth pointing out that Perel'man has shown that on a compact manifold Ricci solitons are always gradient solitons, see [18].

Clearly Einstein metrics are solitons with f being trivial. Another interesting special case occurs when $f = \frac{\lambda}{2}|x|^2$ on \mathbb{R}^n . In this case

$$\text{Hess} f = \lambda g$$

and therefore yields a gradient soliton where the background metric is flat. This example is called a *Gaussian*. Taking a product $N \times \mathbb{R}^k$ with N being Einstein with Einstein constant λ and $f = \frac{\lambda}{2}|x|^2$ on \mathbb{R}^k yields a mixed gradient soliton. We can further take a quotient $N \times_{\Gamma} \mathbb{R}^k$, where Γ acts freely on N and by orthogonal transformations on \mathbb{R}^k (no translational components) to get a flat vector bundle over a base that is Einstein and with $f = \frac{\lambda}{2}d^2$ where d is the distance in the flat fibers to the base.

We say that a gradient soliton is *rigid* if it is of the type $N \times_{\Gamma} \mathbb{R}^k$ just described.

The goal of this paper is to determine when gradient solitons are rigid. For compact manifolds it is easy to see that they are rigid precisely when the scalar curvature is constant see [9]. In fact we can show something a bit more general

Theorem 1.1. *A compact gradient soliton is rigid with trivial f if*

$$\text{Ric}(\nabla f, \nabla f) \leq 0.$$

1991 *Mathematics Subject Classification.* 53C25.

Moreover, in dimensions 2 [11] and 3 [12] all compact solitons are rigid. There are compact shrinking (Kähler) gradient solitons in dimension 4 that do not have constant scalar curvature, the first example was constructed by Koiso [14] see also [3, 23]. It is also not hard to see that, in any dimension, compact steady or expanding solitons are rigid (see [12] and Corollary 3). In fact, at least in the steady gradient soliton case, this seems to go back to Lichnerowicz, see section 3.10 of [2].

In the noncompact case Perel'man has shown that all 3-dimensional shrinking gradient solitons with nonnegative sectional curvature are rigid [19]. However, in higher dimensions, it is less clear how to detect rigidity. In fact there are expanding Ricci solitons with constant scalar curvature that are not rigid in the above sense. These spaces are left invariant metrics on nilpotent groups constructed by Lauret [15] that are not gradient solitons. For other examples of noncompact gradient solitons with large symmetry groups see [3, 4, 10, 13].

Note that if a soliton is rigid, then the “radial” curvatures vanish, i.e.,

$$R(\cdot, \nabla f) \nabla f = 0,$$

and the scalar curvature is constant. Conversely we just saw that constant scalar curvature and radial Ricci flatness: $\text{Ric}(\nabla f, \nabla f) = 0$ each imply rigidity on compact solitons. In the noncompact case we can show

Theorem 1.2. *A shrinking (expanding) gradient soliton*

$$\text{Ric} + \text{Hess}f = \lambda g$$

is rigid if and only if it has constant scalar curvature and is radially flat, i.e., $\text{sec}(E, \nabla f) = 0$.

While radial flatness seems like a strong assumption, there are a number of weaker conditions that guarantee radial flatness.

Proposition 1. *The following conditions for a shrinking (expanding) gradient soliton*

$$\text{Ric} + \text{Hess}f = \lambda g$$

all imply that the metric is radially flat and has constant scalar curvature

- (1) *The scalar curvature is constant and $\text{sec}(E, \nabla f) \geq 0$ ($\text{sec}(E, \nabla f) \leq 0$.)*
- (2) *The scalar curvature is constant and $0 \leq \text{Ric} \leq \lambda g$ ($\lambda g \leq \text{Ric} \leq 0$.)*
- (3) *The curvature tensor is harmonic.*
- (4) *$\text{Ric} \geq 0$ ($\text{Ric} \leq 0$) and $\text{sec}(E, \nabla f) = 0$.*

Given the above theorem it is easy to see that rigid solitons also satisfy these conditions.

Condition 2 is very similar to a statement by Naber, but our proof is quite different. The following result shows that, for shrinking solitons, the scalar curvature condition is in fact redundant. Thus we are offering an alternate proof for part of Naber's result (see [17]).

Lemma 1.3 (Naber). *If M is a shrinking gradient Ricci Soliton with $0 \leq \text{Ric} \leq \lambda g$, then the scalar curvature is constant.*

There is an interesting relationship between this result and Perel'man's classification in dimension 3. The main part of the classification is to show that there are no noncompact shrinking gradient solitons with positive sectional curvature. Perel'man's proof has two parts, first he shows that such a metric has $\text{scal} \leq 2\lambda$

and then he uses this fact, and the Gauss-Bonnet theorem, to arrive at a contradiction. It is a simple algebraic fact that if $\text{sec} \geq 0$ and $\text{scal} \leq 2\lambda$ then $\text{Ric} \leq \lambda$. Therefore, Naber's lemma implies the following gap theorem which generalizes the second part of Perelman's argument to higher dimensions.

Theorem 1.4. *If M^n is a shrinking gradient Ricci soliton with nonnegative sectional curvature and $\text{scal} \leq 2\lambda$ then the universal cover of M is isometric to either \mathbb{R}^n or $S^2 \times \mathbb{R}^{n-2}$.*

The key to most of our proofs rely on a new equation that in a fairly obvious way relates rigidity, radial curvatures, and scalar curvature

$$\nabla_{\nabla f} \text{Ric} + \text{Ric} \circ (\lambda I - \text{Ric}) = R(\cdot, \nabla f) \nabla f + \frac{1}{2} \nabla \cdot \nabla \text{scal}.$$

While we excluded steady solitons from the above result, it wasn't really necessary to do so. In fact it is quite easy to prove something that sounds more general.

Theorem 1.5. *A steady soliton*

$$\text{Ric} + \text{Hess} f = 0$$

whose scalar curvature achieves its minimum is Ricci flat. In particular, steady gradient solitons with constant scalar curvature are Ricci flat.

In the context of condition 3 about harmonicity of the curvature there is a rather interesting connection with gradient solitons. Consider the exterior covariant derivative

$$d^\nabla : \Omega^p(M, TM) \rightarrow \Omega^{p+1}(M, TM)$$

for forms with values in the tangent bundle. The curvature can then be interpreted as the 2-form

$$R(X, Y)Z = ((d^\nabla \circ d^\nabla)(Z))(X, Y)$$

and Bianchi's second identity as $d^\nabla R = 0$. The curvature is harmonic if $d^*R = 0$, where d^* is the adjoint of d^∇ . If we think of Ric as a 1-form with values in TM , then Bianchi's second identity implies

$$d^\nabla \text{Ric} = -d^*R.$$

Thus the curvature tensor is harmonic if and only if the Ricci tensor is closed. This condition has been studied extensively as a generalization of being an Einstein metric (see [1], Chapter 16). It is also easy to see that it implies constant scalar curvature.

Next note that the condition for being a steady gradient soliton is the same as saying that the Ricci tensor is exact

$$\text{Ric} = d^\nabla(-X) = -\nabla X.$$

Since the Ricci tensor is symmetric, this requires that X is locally a gradient field. The general gradient soliton equation

$$\text{Ric} = d^\nabla(-X) + \lambda I$$

then appears to be a simultaneous generalization of being Einstein and exact. Thus Theorem 1.2 implies that rigid gradient solitons are precisely those metrics that satisfy all the generalized Einstein conditions.

Throughout the paper we also establish several other simple results that guarantee rigidity under slightly different assumptions on the curvature and geometry

of the space. We can also use the techniques developed here to obtain some results for solitons with large amounts symmetry, this will be the topic of a forthcoming paper.

2. FORMULAS

In this section we establish the general formulas that will be used to prove the various rigidity results we are after. There are two sets of results. The most general and weakest for Ricci solitons and the more interesting and powerful for gradient solitons.

First we establish a general formula that leads to the Bochner formulas for Killing and gradient fields (see also [21].)

Lemma 2.1. *On a Riemannian manifold*

$$\operatorname{div}(L_X g)(X) = \frac{1}{2}\Delta|X|^2 - |\nabla X|^2 + \operatorname{Ric}(X, X) + D_X \operatorname{div} X$$

When $X = \nabla f$ is a gradient field we have

$$(\operatorname{div} L_X g)(Z) = 2\operatorname{Ric}(Z, X) + 2D_Z \operatorname{div} X$$

or in $(1, 1)$ -tensor notation

$$\operatorname{div} \nabla \nabla f = \operatorname{Ric}(\nabla f) + \nabla \Delta f$$

Proof. We calculate with a frame that is parallel at p

$$\begin{aligned} & \operatorname{div}(L_X g)(X) \\ &= (\nabla_{E_i} L_X g)(E_i, X) \\ &= \nabla_{E_i}(L_X g(E_i, X)) - L_X g(E_i, \nabla_{E_i} X) \\ &= \nabla_{E_i}(g(\nabla_{E_i} X, X) + g(E_i, \nabla_X X)) - g(\nabla_{E_i} X, \nabla_{E_i} X) - g(E_i, \nabla_{\nabla_{E_i} X} X) \\ &= \Delta \frac{1}{2}|X|^2 + \nabla_{E_i} g(E_i, \nabla_X X) - |\nabla X|^2 - g(E_i, \nabla_{\nabla_{E_i} X} X) \\ &= \Delta \frac{1}{2}|X|^2 - |\nabla X|^2 + g(\nabla_{E_i, X}^2 X, E_i) \\ &= \Delta \frac{1}{2}|X|^2 - |\nabla X|^2 + \operatorname{Ric}(X, X) + g(\nabla_{X, E_i}^2 X, E_i) \\ &= \Delta \frac{1}{2}|X|^2 - |\nabla X|^2 + \operatorname{Ric}(X, X) + D_X \operatorname{div} X \end{aligned}$$

And when $Z \rightarrow \nabla_Z X$ is self-adjoint we have

$$\begin{aligned} & (\operatorname{div} L_X g)(Z) \\ &= (\nabla_{E_i} L_X g)(E_i, Z) \\ &= \nabla_{E_i}(L_X g(E_i, Z)) - L_X g(E_i, \nabla_{E_i} Z) \\ &= \nabla_{E_i}(g(\nabla_{E_i} X, Z) + g(E_i, \nabla_Z X)) - g(\nabla_{E_i} X, \nabla_{E_i} Z) - g(E_i, \nabla_{\nabla_{E_i} Z} X) \\ &= \nabla_{E_i}(g(\nabla_{E_i} X, Z) + g(E_i, \nabla_Z X)) - g(\nabla_{E_i} X, \nabla_{E_i} Z) - g(E_i, \nabla_{\nabla_{E_i} Z} X) \\ &= \nabla_{E_i}(g(\nabla_Z X, E_i) + g(E_i, \nabla_{E_i} \nabla_Z X)) - g(\nabla_{E_i} X, \nabla_{E_i} Z) - g(E_i, \nabla_{\nabla_{E_i} Z} X) \\ &= 2g(\nabla_{E_i, Z}^2 X, E_i) \\ &= 2\operatorname{Ric}(Z, X) + 2g(\nabla_{Z, E_i}^2 X, E_i) \\ &= 2\operatorname{Ric}(Z, X) + 2D_Z \operatorname{div} X \end{aligned}$$

□

Corollary 1. *If X is a Killing field, then*

$$\Delta \frac{1}{2} |X|^2 = |\nabla X|^2 - \text{Ric}(X, X)$$

Proof. Use that $L_X g = 0 = \text{div} X$ in the above formula. □

Corollary 2. *If X is a gradient field, then*

$$\Delta \frac{1}{2} |X|^2 = |\nabla X|^2 + D_X \text{div} X + \text{Ric}(X, X)$$

Proof. Let $Z = X$ in the second equation above and equate them to get the formula. □

We are now ready to derive formulas for Ricci solitons

$$\text{Ric} + \frac{1}{2} L_X g = \lambda g$$

Lemma 2.2. *A Ricci soliton satisfies*

$$\frac{1}{2} (\Delta - D_X) |X|^2 = |\nabla X|^2 - \lambda |X|^2$$

Proof. The trace of the soliton equation says that

$$\text{scal} + \text{div} X = n\lambda$$

so

$$D_Z \text{scal} = -D_Z \text{div} X$$

The contracted second Bianchi identity that forms the basis for Einstein's equations says that

$$D_Z \text{scal} = 2 \text{div} \text{Ric}(Z)$$

Using $Z = X$ and the soliton equation then gives

$$\begin{aligned} -D_X \text{div} X &= 2 \text{div} \text{Ric}(X) \\ &= -\text{div}(L_X g)(X) \\ &= -\left(\frac{1}{2} \Delta |X|^2 - |\nabla X|^2 + \text{Ric}(X, X) + D_X \text{div} X \right) \end{aligned}$$

Thus

$$\begin{aligned} \frac{1}{2} \Delta |X|^2 &= |\nabla X|^2 - \text{Ric}(X, X) \\ &= |\nabla X|^2 + \frac{1}{2} (L_X g)(X, X) - \lambda |X|^2 \\ &= |\nabla X|^2 + \frac{1}{2} D_X |X|^2 - \lambda |X|^2 \end{aligned}$$

from which we get the equation. □

We now turn our attention to gradient solitons. In this case we can use $(1, 1)$ -tensors and write the soliton equation as

$$\text{Ric} + \nabla \nabla f = \lambda I$$

or in condensed form

$$\begin{aligned}\operatorname{Ric} + S &= \lambda I, \\ S &= \nabla \nabla f\end{aligned}$$

With this notation we can now state and prove some interesting formulas for the scalar curvature of gradient solitons. The first and last are known (see [7]), while the middle ones seem to be new.

Lemma 2.3. *A gradient soliton satisfies*

$$\nabla \operatorname{scal} = 2\operatorname{Ric}(\nabla f)$$

$$\begin{aligned}\nabla_{\nabla f} S + S \circ (S - \lambda I) &= -R(\cdot, \nabla f) \nabla f - \frac{1}{2} \nabla \cdot \nabla \operatorname{scal}, \\ \nabla_{\nabla f} \operatorname{Ric} + \operatorname{Ric} \circ (\lambda I - \operatorname{Ric}) &= R(\cdot, \nabla f) \nabla f + \frac{1}{2} \nabla \cdot \nabla \operatorname{scal} \\ \frac{1}{2} (\Delta - D_{\nabla f}) \operatorname{scal} &= \frac{1}{2} \Delta_f \operatorname{scal} = \operatorname{tr}(\operatorname{Ric} \circ (\lambda I - \operatorname{Ric}))\end{aligned}$$

Proof. We have the Bochner formula

$$\operatorname{div}(\nabla \nabla f) = \operatorname{Ric}(\nabla f) + \nabla \Delta f$$

The trace of the soliton equation gives

$$\begin{aligned}\operatorname{scal} + \Delta f &= n\lambda, \\ \nabla \operatorname{scal} + \nabla \Delta f &= 0\end{aligned}$$

while the divergence of the soliton equation gave us

$$\operatorname{div} \operatorname{Ric} + \operatorname{div}(\nabla \nabla f) = 0$$

Together this yields

$$\begin{aligned}\nabla \operatorname{scal} &= 2\operatorname{div} \operatorname{Ric} \\ &= -2\operatorname{div}(\nabla \nabla f) \\ &= -2\operatorname{Ric}(\nabla f) - 2\nabla \Delta f \\ &= -2\operatorname{Ric}(\nabla f) + 2\nabla \operatorname{scal}\end{aligned}$$

and hence the first formula.

Using this one can immediately find a formula for the Laplacian of the scalar curvature. However our goal is the establish the second set of formulas. The last formula is then obtained by taking traces.

We use the equation

$$R(E, \nabla f) \nabla f = \nabla_{E, \nabla f}^2 \nabla f - \nabla_{\nabla f, E}^2 \nabla f$$

The second term on the right

$$\nabla_{\nabla f, E}^2 \nabla f = (\nabla_{\nabla f} S)(E)$$

while the first can be calculated

$$\begin{aligned}
\nabla_{E, \nabla f}^2 \nabla f &= -(\nabla_E \text{Ric})(\nabla f) \\
&= -\nabla_E \text{Ric}(\nabla f) + \text{Ric}(\nabla_E \nabla f) \\
&= -\frac{1}{2} \nabla_E \nabla \text{scal} + \text{Ric} \circ S(E) \\
&= -\frac{1}{2} \nabla_E \nabla \text{scal} + (\lambda I - S) \circ S(E) \\
&= -\frac{1}{2} \nabla_E \nabla \text{scal} + \text{Ric} \circ (\lambda I - \text{Ric})
\end{aligned}$$

This yields the set of formulas in the middle.

Taking traces in

$$\nabla_{\nabla f} \text{Ric} + \text{Ric} \circ (\lambda I - \text{Ric}) = R(E, \nabla f) \nabla f + \frac{1}{2} \nabla_E \nabla \text{scal}$$

yields

$$\nabla_{\nabla f} \text{scal} + \text{tr}(\text{Ric} \circ (\lambda I - \text{Ric})) = \text{Ric}(\nabla f, \nabla f) + \frac{1}{2} \Delta \text{scal}$$

Since

$$\text{Ric}(\nabla f, \nabla f) = \frac{1}{2} D_{\nabla f} \text{scal}$$

we immediately get the last equation. \square

Note that if λ_i are the eigenvalues of the Ricci tensor then the last equation can be rewritten in several useful ways

$$\begin{aligned}
\frac{1}{2} \Delta_f \text{scal} &= \text{tr}(\text{Ric} \circ (\lambda I - \text{Ric})) \\
&= \sum \lambda_i (\lambda - \lambda_i) \\
&= -|\text{Ric}|^2 + \lambda \text{scal} \\
&= -\left| \text{Ric} - \frac{1}{n} \text{scal} g \right|^2 + \text{scal} \left(\lambda - \frac{1}{n} \text{scal} \right)
\end{aligned}$$

3. RIGIDITY CHARACTERIZATION

We start with a motivational appetizer on rigidity of gradient solitons.

Proposition 2. *A gradient soliton which is Einstein, either has $\text{Hess} f = 0$ or is a Gaussian.*

Proof. Assume that

$$\mu g + \text{Hess} f = \lambda g.$$

If $\mu = \lambda$, then the Hessian vanishes. Otherwise we have that the Hessian is proportional to g . Multiplying f by a constant then leads us to a situation where

$$\text{Hess} f = g.$$

This shows that f is a proper strictly convex function. By adding a suitable constant to f we also see that $r = \sqrt{f}$ is a distance function from the unique minimum of f . It is now easy to see that the radial curvatures vanish and then that the space is flat (see also [20]) \square

Next we dispense with rigidity for compact solitons.

Theorem 3.1. *A compact Ricci soliton with*

$$\operatorname{Ric}(X, X) \leq 0$$

is Einstein with Einstein constant λ . In particular, compact gradient solitons with constant scalar curvature are Einstein.

Proof. We have a Ricci soliton

$$\operatorname{Ric} + L_X g = \lambda g.$$

The Laplacian of X then satisfies

$$\begin{aligned} \Delta \frac{1}{2} |X|^2 &= |\nabla X|^2 - \operatorname{Ric}(X, X) \\ &\geq 0 \end{aligned}$$

The divergence theorem then shows that ∇X vanishes. In particular $L_X g = 0$.

The second part is a simple consequence of having $X = \nabla f$ and the equation

$$D_{\nabla f} \operatorname{scal} = 2\operatorname{Ric}(\nabla f, \nabla f).$$

□

We also note that, when the Ricci tensor has a definite sign, having zero radial Ricci curvature is equivalent to having constant scalar curvature. In particular this implies the equivalence of condition (4) in Proposition 1.

Proposition 3. *A gradient soliton with nonnegative (or nonpositive) Ricci curvature has constant scalar curvature if and only if $\operatorname{Ric}(\nabla f, \nabla f) = 0$.*

Proof. We know from elementary linear algebra that, for a nonnegative (or non-positive) definite, self-adjoint operator T ,

$$\langle Tv, v \rangle = 0 \quad \Rightarrow \quad Tv = 0.$$

So the proposition follows easily by taking T to be the $(1, 1)$ -Ricci tensor and the fact that $\nabla \operatorname{scal} = 2\operatorname{Ric}(\nabla f)$ for a gradient soliton. □

Steady solitons are also easy to deal with

Proposition 4. *A steady gradient soliton with constant scalar curvature is Ricci flat. Moreover, if f is not constant then it is a product of a Ricci flat manifold with \mathbb{R} .*

Proof. First we note that

$$\begin{aligned} 0 &= \frac{1}{2} \Delta_f \operatorname{scal} \\ &= - \left| \operatorname{Ric} - \frac{1}{n} \operatorname{scal} g \right|^2 + \operatorname{scal} \left(\lambda - \frac{1}{n} \operatorname{scal} \right) \\ &= - \left| \operatorname{Ric} - \frac{1}{n} \operatorname{scal} g \right|^2 - \frac{1}{n} \operatorname{scal}^2 \\ &\leq 0 \end{aligned}$$

Thus $\operatorname{scal} = 0$ and $\operatorname{Ric} = 0$. This shows that $\operatorname{Hess} f = 0$. Thus f is either constant or the manifold splits along the gradient of f . □

This partly motivates our next result.

Proposition 5. *Assume that we have a gradient soliton*

$$\text{Ric} + \text{Hess}f = \lambda g$$

with constant scalar curvature and $\lambda \neq 0$. When $\lambda > 0$ we have $0 \leq \text{scal} \leq n\lambda$. When $\lambda < 0$ we have $n\lambda \leq \text{scal} \leq 0$. In either case the metric is Einstein when the scalar curvature equals either of the extreme values.

Proof. Again we have that

$$0 = \frac{1}{2}\Delta_f \text{scal} = - \left| \text{Ric} - \frac{1}{n}\text{scal}g \right|^2 + \text{scal} \left(\lambda - \frac{1}{n}\text{scal} \right)$$

showing that

$$0 \leq \left| \text{Ric} - \frac{1}{n}\text{scal}g \right|^2 = \text{scal} \left(\lambda - \frac{1}{n}\text{scal} \right)$$

Thus $\text{scal} \in [0, n\lambda]$ if the soliton is shrinking and the metric is Einstein if the scalar curvature takes on either of the boundary values. A similar analysis holds in the expanding case. \square

Before proving the main characterization we study the conditions that guarantee radial flatness.

Proposition 6. *The following conditions for a shrinking (expanding) gradient soliton*

$$\text{Ric} + \text{Hess}f = \lambda g$$

all imply that it is radially flat.

- (1) *The scalar curvature is constant and $\sec(E, \nabla f) \geq 0$ ($\sec(E, \nabla f) \leq 0$).*
- (2) *The scalar curvature is constant and $0 \leq \text{Ric} \leq \lambda g$ ($\lambda g \leq \text{Ric} \leq 0$).*
- (3) *The curvature tensor is harmonic.*

Proof. 1: Use the equations

$$\begin{aligned} 0 &= \frac{1}{2}\nabla_{\nabla f} \text{scal} = \text{Ric}(\nabla f, \nabla f) \\ &= \sum g(R(E_i, \nabla f)\nabla f, E_i) \end{aligned}$$

to see that $g(R(E_i, \nabla f)\nabla f, E_i) = 0$ if the radial curvatures are always nonnegative (nonpositive).

2: First observe that

$$0 = \frac{1}{2}\Delta_f \text{scal} = \text{tr}(\text{Ric} \circ (\lambda I - \text{Ric}))$$

The assumptions on the Ricci curvature imply that $\text{Ric} \circ (\lambda I - \text{Ric})$ is a nonnegative operator. Thus

$$\text{Ric} \circ (\lambda I - \text{Ric}) = 0.$$

This shows that the only possible eigenvalues for Ric and $\nabla \nabla f$ are 0 and λ .

To establish radial flatness we then use that the formula

$$\nabla_{\nabla f} \text{Ric} + \text{Ric} \circ (\lambda I - \text{Ric}) = R(\cdot, \nabla f)\nabla f + \frac{1}{2}\nabla \cdot \nabla \text{scal}$$

is reduced to

$$\begin{aligned} R(\cdot, \nabla f)\nabla f &= \nabla_{\nabla f} \text{Ric} \\ &= -\nabla_{\nabla f, \cdot}^2 \nabla f \end{aligned}$$

Next pick a field E such that $\nabla_E \nabla f = 0$, then

$$\begin{aligned} g(\nabla_{\nabla f, E}^2 \nabla f, E) &= g(\nabla_{\nabla f} \nabla_E \nabla f, E) - g(\nabla_{\nabla_{\nabla f} E} \nabla f, E) \\ &= -g(\nabla_E \nabla f, \nabla_{\nabla f} E) \\ &= 0 \end{aligned}$$

and finally when $\nabla_E \nabla f = \lambda E$

$$\begin{aligned} g(\nabla_{\nabla f, E}^2 \nabla f, E) &= g(\nabla_{\nabla f} \nabla_E \nabla f, E) - g(\nabla_{\nabla_{\nabla f} E} \nabla f, E) \\ &= \lambda g(\nabla_{\nabla f} E, E) - g(\nabla_E \nabla f, \nabla_{\nabla f} E) \\ &= \lambda g(\nabla_{\nabla f} E, E) - \lambda g(E, \nabla_{\nabla f} E) \\ &= 0. \end{aligned}$$

Thus $g(R(E, \nabla f) \nabla f, E) = 0$ for all eigenfields. This shows that the metric is radially flat.

3: Finally use the soliton equation to see that

$$(\nabla_X \text{Ric})(Y, Z) - (\nabla_Y \text{Ric})(X, Z) = -g(R(X, Y) \nabla f, Z).$$

From the 2nd Bianchi identity we also get that

$$(\nabla_X \text{Ric})(Y, Z) - (\nabla_Y \text{Ric})(X, Z) = \text{div} R(X, Y, Z) = 0$$

since the curvature is harmonic. Thus $R(X, Y) \nabla f = 0$. In particular $\text{sec}(E, \nabla f) = 0$. \square

We now turn our attention to the main theorem. To prepare the way we show.

Proposition 7. *Assume that we have a gradient soliton*

$$\text{Ric} + \text{Hess} f = \lambda g$$

with constant scalar curvature, $\lambda \neq 0$ and a nontrivial f . For a suitable constant α

$$f + \alpha = \frac{\lambda}{2} r^2$$

where r is a smooth function whenever $\nabla f \neq 0$ and satisfies

$$|\nabla r| = 1.$$

Proof. Observe that

$$\begin{aligned} \frac{1}{2} \nabla (\text{scal} + |\nabla f|^2) &= \text{Ric}(\nabla f) + \nabla_{\nabla f} \nabla f \\ &= \lambda \nabla f \end{aligned}$$

which shows

$$\text{scal} + |\nabla f|^2 - 2\lambda f = \text{const}$$

By adding a suitable constant to f we can then assume that

$$|\nabla f|^2 = 2\lambda f.$$

Thus f has the same sign as λ and the same zero locus as its gradient. If we define r such that

$$f = \frac{\lambda}{2} r^2$$

then

$$\nabla f = \lambda r \nabla r$$

and

$$\begin{aligned} 2\lambda f &= |\nabla f|^2 \\ &= \lambda^2 r^2 |\nabla r|^2 \\ &= 2\lambda f |\nabla r|^2 \end{aligned}$$

□

This allows us to establish our characterization of rigid gradient solitons.

Theorem 3.2. *A gradient soliton*

$$\text{Ric} + \text{Hess} f = \lambda g$$

is rigid if it is radially flat and has constant scalar curvature.

Proof. We consider the case where $\lambda > 0$ as the other case is similar aside from some sign changes.

Using the condensed version of the soliton equation

$$\begin{aligned} \text{Ric} + S &= \lambda I, \\ S &= \nabla \nabla f \end{aligned}$$

we have

$$\begin{aligned} \nabla_{\nabla f} S + S \circ (S - \lambda I) &= 0, \\ \nabla_{\nabla f} \text{Ric} + \text{Ric} \circ (\lambda I - \text{Ric}) &= 0 \end{aligned}$$

Assume that $f = \frac{\lambda}{2} r^2$ where r is a nonnegative distance function. The minimum set for f

$$N = \{x : f(x) = 0\}$$

is also characterized as

$$N = \{x \in M : \nabla f(x) = 0\}$$

This shows that $S \circ (S - \lambda I) = 0$ on N .

When $r > 0$ we note that the smallest eigenvalue for S is always absolutely continuous and therefore satisfies the differential equation

$$D_{\nabla f} \mu_{\min} = \mu_{\min} (\lambda - \mu_{\min}).$$

We claim that $\mu_{\min} \geq 0$. Using $r > 0$ as an independent coordinate and $\nabla f = \lambda r \nabla r$ yields

$$\partial_r \mu_{\min} = \frac{1}{\lambda r} \mu_{\min} (\lambda - \mu_{\min})$$

This equation can be solved by separation of variables. In particular, $\mu_{\min} \rightarrow -\infty$ in finite time provided $\mu_{\min} < 0$ somewhere. This contradicts smoothness of f . Thus we can conclude that $\mu_{\min} \geq 0$ and hence that f is convex.

Now that we know f is convex the minimum set N must be totally convex. We also know that on N the eigenvalues of $\nabla \nabla f$ can only be 0 and λ . Thus their multiplicities are constant. Using that the rank of $\nabla \nabla f$ is constant we see that N is a submanifold whose tangent space is given by $\ker(\nabla \nabla f)$. This in turn shows that N is a totally geodesic submanifold.

Note that when $\lambda > 0$ the minimum set N is in fact compact as it must be an Einstein manifold with Einstein constant λ .

The normal exponential map

$$\exp : \nu(N) \rightarrow M$$

follows the integral curves for ∇f or ∇r and is therefore a diffeomorphism.

Using the fundamental equations (see [20]) we see that the metric is completely determined by the fact that it is radially flat and that N is totally geodesic. From this it follows that the bundle is flat and hence of the type $N \times_{\Gamma} \mathbb{R}^k$.

Alternately note that radial flatness shows that all Jacobi fields along geodesics tangent to ∇f must be of the form

$$J = E + tF$$

where E and F are parallel. This also yields the desired vector bundle structure. \square

4. OTHER RESULTS

In this section we discuss some further applications of the formulas derived above. First we recall some technical tools. We will use the following notation.

$$\Delta_X = \Delta - D_X$$

Recall the maximum principle for elliptic PDE's.

Theorem 4.1 (Maximum Principle). *If u is a real valued function with $\Delta_X(u) \geq 0$ then u is constant in a neighborhood of any local maximum.*

The first lemma follows from Lemma 2.2.

Lemma 4.2. *If M is a complete expanding or steady Ricci soliton then*

$$\Delta_X |X|^2 \geq 0.$$

Moreover, $\Delta_X |X|^2 = 0$ if and only if M is Einstein.

Proof. This follows directly from the formula $\frac{1}{2}\Delta_X |X|^2 = |\nabla X|^2 - \lambda |X|^2$. \square

Applying the maximum principle then shows that $|X|$ can not achieve its maximum without being trivial.

Theorem 4.3. *If M is a complete expanding or steady Ricci soliton and $|X|$ achieves its maximum then M is Einstein.*

Note that this clearly implies the following result for compact steady and expanding solitons mentioned in the introduction.

Corollary 3. *Compact expanding or steady Ricci solitons are Einstein.*

When we have a gradient soliton we use the notation $\Delta_X = \Delta_f$. From Lemma 2.3 we also have the following inequality

Lemma 4.4. *If M is a steady gradient soliton or an expanding gradient soliton with nonnegative scalar curvature, then*

$$\Delta_f(\text{scal}) \leq 0.$$

Moreover, $\Delta_f(\text{scal}) = 0$ if and only if M is Ricci flat. In particular, the only expanding gradient soliton with nonnegative scalar curvature and $\Delta_f(\text{scal}) = 0$ is the Gaussian.

Proof. This follows easily from the equation

$$\frac{1}{2}\Delta_f \text{scal} = -|\text{Ric}|^2 + \lambda \text{scal}.$$

That a Ricci flat expanding soliton must be a Gaussian is just Proposition 2. \square

Now from the maximum principle we have that the scalar curvature cannot have a minimum.

Theorem 4.5. (1) *A steady gradient soliton whose scalar curvature achieves its minimum is Ricci flat.*
 (2) *An expanding gradient soliton with nonnegative scalar curvature achieving its minimum is a Gaussian.*

For gradient solitons there is a naturally associated measure $dm = e^{-f} d\text{vol}_g$ which makes the operator Δ_f self-adjoint. Namely the following identity holds for compactly supported functions.

$$\int_M \Delta_f(\phi)\psi dm = - \int_M \langle \nabla\phi, \nabla\psi \rangle dm = \int_M \phi \Delta_f(\psi) dm.$$

The measure dm also plays an important role in Perel'man's entropy formulas for the Ricci flow [18]. In [24], Yau proves that on a complete Riemannian manifold any L^α , positive, subharmonic function is constant. The argument depends solely on using integration by parts and picking a clever test function ϕ . Therefore, the argument completely generalizes to the measure dm and operator Δ_f . Specifically the following L^α Liouville theorem holds.

Theorem 4.6 (Yau). *Any nonnegative real valued function u with $\Delta_f(u)(x) \geq 0$ which satisfies the condition*

$$(4.1) \quad \lim_{r \rightarrow \infty} \left(\frac{1}{r^2} \int_{B(p,r)} u^\alpha dm \right) = 0$$

for some $\alpha > 1$ is constant.

Define $\Omega_{u,C} = \{x : u(x) \geq C\}$. If we only have a bound on the f -Laplacian on $\Omega_{u,C}$ then we can apply the L^α Liouville theorem to prove the following corollary.

Corollary 4. *If $\Delta_f(u)(x) \geq 0$ for all $x \in \Omega_{u,C}$ and u satisfies (4.1) then u is either constant or $u \leq C$.*

Proof. Apply Theorem 4.6 to the function $(u-C)_+ = \max\{u-C, 0\}$. Then $(u-C)_+$ is constant which implies either $u \leq C$ or u is constant. \square

One can also derive upper bounds on the growth of the measure dm from the inequality $\text{Ric} + \text{Hess}f \geq \lambda g$ see [16, 22]. In particular, when $\lambda > 0$, the measure is bounded above by a Gaussian measure. Combining this estimate with the L^α maximum principle gives the following strong Liouville theorem for shrinking gradient Ricci solitons.

Corollary 5. [22] *If M is a complete manifold satisfying*

$$\text{Ric} + \text{Hess}f \geq \lambda g$$

for $\lambda > 0$ and u is a real valued function such that $\Delta_f(u) \geq 0$ and $u(x) \leq Ke^{\beta d(p,x)^2}$ for some $\beta < \lambda$ then u is constant.

A similar result, under the additional assumption that Ric is bounded above, is proven by Naber [17]. In fact, one can see immediately from the equation

$$\Delta_f(\text{scal}) = \sum \lambda_i(\lambda - \lambda_i)$$

that if $0 \leq \text{Ric} \leq \lambda$ for a shrinking soliton then scal is bounded, nonnegative, and has $\Delta_f(\text{scal}) \geq 0$. Therefore it is constant and we have Lemma 1.3. Using the Liouville theorem the following improvement of Proposition 5 is also true for shrinking gradient solitons.

Theorem 4.7. *If scal is bounded, then*

$$0 \leq \inf_M \text{scal} \leq n\lambda.$$

Moreover, if $\text{scal} \geq n\lambda$, then M is Einstein.

Proof. First suppose that $\text{scal} \geq n\lambda$. By the Cauchy-Schwarz inequality

$$\begin{aligned} \Delta_f(\text{scal}) &= -|\text{Ric}|^2 + \lambda \text{scal} \\ &\leq -\frac{\text{scal}^2}{n} + \lambda \text{scal} \\ &\leq \text{scal} \left(\lambda - \frac{\text{scal}}{n} \right) \end{aligned}$$

So that $\Delta_f(\text{scal}) \leq 0$. Let K be the upper bound on scal then the function $u = K - \text{scal}$ is bounded, nonnegative, and has $\Delta_f(u) \geq 0$. So by Corollary 5 scal is constant and thus must be Einstein.

To see the other inequality consider that on $\Omega_0 = \{x : \text{scal}(x) \leq 0\}$, $\Delta_f(\text{scal}) \leq 0$, so applying Corollary 4 to $-\text{scal}$ gives the result. \square

For steady and expanding gradient solitons we can also apply the L^α Liouville theorem to the equation, $\Delta_f(|\nabla f|^2) \geq 0$.

Theorem 4.8. *Let $\alpha > 2$. If M is a steady or expanding soliton with*

$$(4.2) \quad \limsup_{r \rightarrow \infty} \frac{1}{r^2} \int_{B(p,r)} |\nabla f|^\alpha e^{-f} d\text{vol}_g = 0.$$

then M is Einstein.

We think of Theorem 4.8 as a gap theorem for the quantity $\int_{B(p,r)} |\nabla f|^\alpha e^{-f} d\text{vol}_g$ since, if M is Einstein, the quantity is zero.

For steady solitons $\text{scal} + |\nabla f|^2$ is constant so if the scalar curvature is bounded then so is $|\nabla f|$ and (4.2) is equivalent to the measure dm growing sub-quadratically. Therefore, we have the following corollary.

Corollary 6. *If M is a steady Ricci soliton with bounded scalar curvature and*

$$\lim_{r \rightarrow \infty} \frac{1}{r^2} \int_{B(p,r)} e^{-f} d\text{vol}_g = 0$$

Then M is Ricci flat.

We note the relation of this result to the theorem proved by the second author and Wei that if $\text{Ric} + \text{Hess}f \geq 0$ and f is bounded then the growth of $e^{-f} d\text{vol}_g$ is at least linear [22]. Since Ricci flat manifolds have at least linear volume growth Corollary 6 implies that steady Ricci solitons with bounded scalar curvature also have at least linear dm -volume growth. There are Ricci flat manifolds with linear volume growth so Corollary 6 can be viewed as a gap theorem for the growth of dm on gradient steady solitons.

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